## 熱帯の強い降雨時に Swarm 衛星が東南アジア上空で観測した磁場変動

# 家森 俊彦 [1]; 青山 忠司 [2]; 山田 彬泰 [3]; Hozumi Kornyanat[4]; 中西 邦仁 [5]; 横山 佳弘 [6]; 佐納 康治 [7]; 小田木 洋子 [8]; Jarupongsakul Thanawat[9]; Pangsapa Vijak[10]

[1] 京大; [2] 京大・理; [3] 京大・理; [4] NICT; [5] 京都大学理学研究科; [6] 京大理; [7] 朝日大・経営; [8] 京大・理・地磁気センター; [9] チュラロンコン大・理; [10] チュラロンコン大・理

## Magnetic variations observed by the Swarm satellites over south-east Asia during strong tropical rainfall

# Toshihiko Iyemori[1]; Tadashi Aoyama[2]; Akiyasu Yamada[3]; Kornyanat Hozumi[4]; Kunihito Nakanishi[5]; Yoshihiro Yokoyama[6]; Yasuharu Sano[7]; Yoko Odagi[8]; Thanawat Jarupongsakul[9]; Vijak Pangsapa[10]
[1] Kyoto Univ.; [2] Graduate School of Science, Kyoto Univ.; [3] Faculty of Science, Kyoto University; [4] NICT; [5] Graduate

School of Science, Kyoto Univ.; [2] Graduate School of Science, Kyoto Univ.; [3] Faculty of Science, Kyoto University; [4] NIC1; [5] Graduate School of Science, Kyoto Univ.; [7] Asahi Univ.; [8] WDC for Geomagnetism, Kyoto Univ.; [9] Faculty of Science, Chulalongkorn Univ.; [10] Chulalongkorn Univ.

Cumulative convection is expected to generate acoustic mode atmospheric waves, and they are expected to generate Magnetic Ripples. They are observed as a small scale magnetic variation. That is, their typical amplitude is less than a few nT with period around 10-30 seconds. They are observed by low-altitude satellites almost always along the orbit in low and mid latitudes. From the Swarm satellite observation, it was confirmed that they are spatial structure of short scale field-aligned currents<sup>(1)</sup>. Various case studies and statistical analyses strongly suggest that the main source is the cumulative convection<sup>(2),(3)</sup> in lower atmosphere. However, it is still not yet very clear probably because the cumulative convection exists everywhere. To show the generation process of magnetic ripples more directly, we have been making geomagnetic and micro-barometric observations, GPS-TEC and meteorological observations such as rain-fall, wind velocity, temperature etc. in Phimai, north-east of Thailand and in Nakanoshima Island South-West Japan. In this paper, we show the results of Swarm satellite observation over South-East Asia, GPS-TEC and ground observations during conjunction events. The polar orbits of Swarm-A and &#8211;C satellites are shifted about 1.4 degree in longitude, i.e., Swarm-C flies about 1.4 degree East of Swarm-A with around 10 seconds delay, and hence, they fly parallel with maximum distance about 140 km near the equator.

- (1) Iyemori et al. (2015), Geophys. Res. Lett., 41, doi:10.1002/2014GL062555.
- (2) Nakanishi et al. (2014), Earth Planets Space, 66:40, doi:10.1186/1880-5981-66-40.
- (3) Aoyama et al. (2017), Earth, Planets and Space, 69:89, DOI 10.1186/s40623-017-0679-2.

低高度極軌道衛星が中低緯度でほぼ常に観測する 磁気リップル と名付けられた小振幅の短周期磁場変動は、微細な沿磁力線電流の空間構造であることが、2013 年 11 月に打ち上げられた ESA の Swarm-A,-B,-C 3 機編隊による精密磁場観測データの解析から明らかになった (1)。振幅の地理的、季節的分布や緯度分布、Local Time 依存性の統計的解析や、全球雲画像、台風との対応など、様々な状況証拠から、下層大気中の積雲対流から放射される音波モードの大気重力波動による電離層でのダイナモ作用が原因であると推測される (2),(3)。しかし、どの積雲対流が発生源となっているのかなど、より直接的な対応をつけることはできていない。そこで、この発表では、熱帯域でのスコールに伴うような急な激しい降雨現象が発生したときに上空を経度方向に約 1.4 度平行して飛翔した Swarm-A と Swarm-C の磁場観測データ、タイ東北部 Phimai やトカラ中之島で同時観測した地磁気、微気圧、降雨、風速および GPS-TEC データなどを相互に比較した結果を報告する。

- (1) Iyemori et al. (2015), Geophys. Res. Lett., 41, doi:10.1002/2014GL062555.
- (2) Nakanishi et al. (2014), Earth Planets Space, 66:40, doi:10.1186/1880-5981-66-40.
- (3) Aoyama et al. (2017), Earth, Planets and Space, 69:89, DOI 10.1186/s40623-017-0679-2.